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## Gas phase ion-molecule reactions of monogermane with carbon oxides and ethylene: formation of germanium–carbon bonds

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### Abstract

The gas phase ion-molecule reactions of  $\text{GeH}_4$  with some carbon-containing compounds ( $\text{CO}$ ,  $\text{CO}_2$ , and  $\text{C}_2\text{H}_4$ ) have been studied by high pressure mass spectrometry and Fourier transform mass spectrometry. The effects of the total pressure and of the relative concentrations of the reagent gases on the ion pattern are reported. In the presence of  $\text{CO}$  and  $\text{CO}_2$ ,  $\text{GeH}_n\text{CO}^+$  or  $\text{GeH}_n\text{CO}_2^+$ ,  $\text{GeH}_n\text{O}^+$ , and  $\text{GeH}_n\text{C}^+$  ions are observed, all of which show very low abundances. In contrast, condensation processes of  $\text{GeH}_4$  with  $\text{C}_2\text{H}_4$  give  $\text{GeC}_n\text{H}_m^+$  ( $n = 1-4$ ) species, in very high yield for  $n = 1, 2$  but lower for  $n = 3, 4$ . For all three systems, reaction mechanisms are suggested and are discussed in relation to the preparation of amorphous materials containing germanium carbides for photovoltaic applications.

### Introduction

The recent discovery that amorphous silicon can be doped has opened up important possibilities, both economic and technical, for photovoltaic cell industry. Amorphous silicon can be obtained in very high purity degree by simple procedures since it is prepared from gaseous reactants from which impurities are easily removed. Even though the amorphous silicon has a higher optical gap than the crystalline one (1.55 eV vs. 1.1 eV) [1], a great deal of the current research in photovoltaic technology is devoted to the amorphous form [2], because of its lower cost. Amorphous layers are generally prepared by deposition methods such as chemical vapour deposition (CVD) [1–5] and glow discharge (GD) [6], which give good results.

In previous studies [7], it was observed that the solid product obtained from volatile hydrides ( $\text{SiH}_4$ ,  $\text{GeH}_4$ ) by radiolytic fragmentation shows a similar composition to that of the analogous materials prepared by GD. The main advantage of the radiolytic method compared with both CVD and GD is that the radical reaction mechanisms in the gas phase can be interpreted. Moreover, the ionic reactions involved in these processes can be studied by chemical ionization mass spectrometry (CIMS) and Fourier transform mass spectrometry (FTMS) in self-condensation experiments [8]. The understanding of these mechanisms can lead to an a priori design of the structural characteristics of the desired solid with obvious economic and technical advantages.

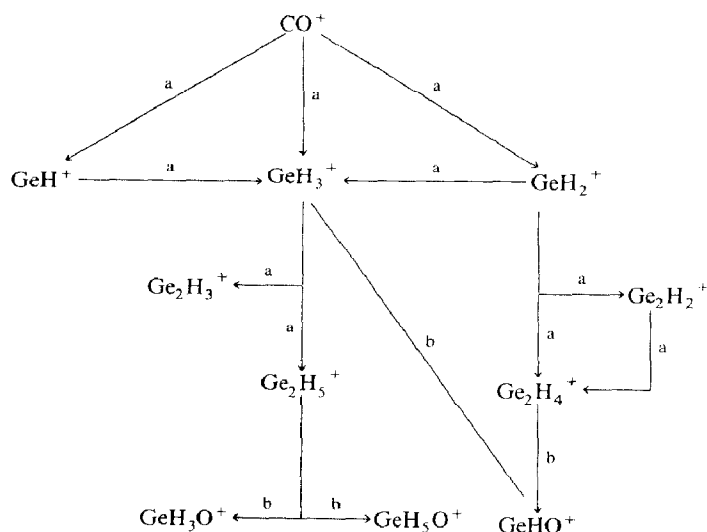
We focused our attention mainly on a study of the preparation and the properties of amorphous germanium, since amorphous silicon had already been well studied by many research groups [9] and also because use of germanium in the photovoltaic industry has very interesting possibilities. Germanium thin layers do not seem suitable to be used in solar cell productions because of their relatively low photoconductivity, but theoretical considerations indicate that germanium carbides ( $\text{Ge}_x\text{C}_{1-x}$ ) have good photovoltaic properties since they show high conversion efficiency [10].

In this paper we report the results of a study of the gas phase ion-molecule reactions of  $\text{GeH}_4$  in the presence of carbon-containing compounds as monitored by Fourier transform mass spectrometry (FTMS) and high pressure mass spectrometry (CIMS). The reactions of  $\text{GeH}_4$  with different carbon-containing molecules,  $\text{CO}$ ,  $\text{CO}_2$  and  $\text{C}_2\text{H}_4$ , together with the reaction mechanisms, are described and discussed with the ultimate aim of identifying the experimental conditions leading to amorphous germanium carbides of the desired composition.

## Results and discussion

### *Ion-molecule reactions in $\text{GeH}_4/\text{CO}$ mixtures*

A 20 eV electron beam on  $\text{GeH}_4/\text{CO}$  mixture (1/25) at the pressure of  $5 \times 10^{-7}$



Scheme 1

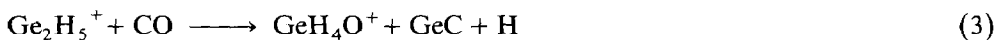
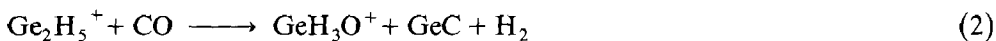
torr in the FTMS cell produces essentially the  $\text{CO}^+$  ion together with appreciable amounts of the primary germane ion family  $\text{GeH}_n^+$  ( $n = 0 - 3$ ). The main reactions, as revealed by double resonance experiments, are described in Scheme 1. Only ionic reagents/products are reported, the neutral reagent molecules being  $\text{GeH}_4$  (a) or  $\text{CO}$  (b). Owing to the higher ionization potential of carbon monoxide,  $14.013 \pm 0.004$  eV [11], compared with that of monogermene, 10.5 eV [12], collisions between  $\text{CO}^+$  and  $\text{GeH}_4$  lead to disappearance of  $\text{CO}^+$  in less than 100 ms, with predominant formation of  $\text{GeH}_3^+$  (ca. 65%),  $\text{GeH}_2^+$  (ca. 20%), and  $\text{GeH}^+$  (ca. 15%) ions (reaction 1).



The  $\text{GeH}_3^+$  ion, which shows the highest abundance, is also formed by reactions of neutral  $\text{GeH}_4$  with  $\text{GeH}_2^+$ , by hydride abstraction, and with  $\text{GeH}^+$  within the same reaction time. Moreover, reaction 1 is the most favoured when  $n = 3$ , as it is exothermic by more than 100 kcal/mol ( $\text{HCO}$ , neutral fragment) compared with ca. 90 kcal/mol and 50 kcal/mol when  $n = 1$  ( $\text{HCO}$  and  $\text{H}_2$ , neutral fragments), and  $n = 2$  ( $\text{H}_2\text{CO}$ , neutral fragment), respectively.

The  $\text{GeH}_3^+$  species react with  $\text{GeH}_4$  in accord with the normal ion pattern described previously [8], yielding essentially  $\text{Ge}_2\text{H}_5^+$  and  $\text{Ge}_2\text{H}_3^+$  (> 82%), while  $\text{GeH}_2^+$  gives minor amounts of  $\text{Ge}_2\text{H}_4^+$  and  $\text{Ge}_2\text{H}_2^+$  (< 18%). These ions lose progressively their hydrogens by collisions with  $\text{CO}$ , leading eventually to  $\text{Ge}_2^+$ . Alternatively,  $\text{Ge}_2\text{H}_5^+$  reacts with a  $\text{CO}$  molecule to yield  $\text{GeH}_5\text{O}^+$ , together with minor quantities of  $\text{GeH}_3\text{O}^+$ , in 200–500 ms reaction time. In addition, the  $\text{GeHO}^+$  ion is observed, which is formed by comparatively slow condensation reactions of  $\text{Ge}_2\text{H}_4^+$  or  $\text{GeH}_3^+$  and  $\text{CO}$ . It should be pointed out that formation of the  $\text{GeH}_n\text{O}^+$  ( $n = 1, 3, 5$ ) ions, which may be regarded as the germanium-containing counterpart of protonated carbon monoxide ( $n = 1$ ), formaldehyde ( $n = 3$ ), and methanol ( $n = 5$ ), respectively, proceeds via complex mechanisms which necessarily involve formation of a Ge–O covalent bond between the germane ions and  $\text{CO}$  and extensive rearrangement of the hydrogen atoms in the resulting adducts.

It is noteworthy that  $\text{Ge}_2\text{H}_5^+$  reacts with neutral  $\text{CO}$  to give only  $\text{GeH}_n\text{O}^+$  ions when  $n$  is an odd number ( $n = 3, 5$ ), and neither  $\text{GeH}_4\text{O}^+$  nor  $\text{GeH}_2\text{O}^+$  species are ever observed in FTMS experiments. The lack of these latter ions is likely to be due to the energetics of their formation reactions, but insufficient thermochemical data are available for calculating the enthalpies of such reactions. However, reaction 2, which has been observed, and reaction 3, which does not occur, in which  $\text{GeH}_3\text{O}^+$  and  $\text{GeH}_4\text{O}^+$  are formed, respectively, differ by only one neutral fragment.



A hydrogen molecule is a neutral fragment in reaction 2, while a hydrogen atom would be lost in reaction 3, and this increases the reaction enthalpy by more than 50 kcal/mol. Thus, the absence of  $\text{GeH}_n\text{O}^+$  ( $n = 2, 4$ ) ions does not imply that they have an higher heat of formation than the corresponding species with  $n = 3, 5$ , but could be due only to the different neutral fragments lost in their formation reactions.

Table 1

Relative abundances of significant ions in the mass spectra of GeH<sub>4</sub>(9.83%)/CO mixtures at various pressures <sup>a</sup>

| Ions  | Pressures (torr) |      |      |
|---|------------------|------|------|
|   | 0.3              | 0.1  | 0.03 |
| Ge <sub>2</sub> H <sub>5</sub> <sup>+</sup> | 0.1              | –    | –    |
| Ge <sub>2</sub> H <sub>4</sub> <sup>+</sup> | 0.3              | 0.2  | 0.1  |
| Ge <sub>2</sub> H <sub>3</sub> <sup>+</sup> | 0.2              | 0.1  | 0.05 |
| Ge <sub>2</sub> H <sub>2</sub> <sup>+</sup> | 0.4              | 0.2  | 0.1  |
| Ge <sub>2</sub> H <sup>+</sup>              | 0.02             | 0.01 | –    |
| GeH <sub>3</sub> CO <sup>+</sup>            | 0.1              | –    | –    |
| GeHCO <sup>+</sup>                          | 0.2              | 0.2  | 0.2  |
| GeCO <sup>+</sup>                           | 0.6              | 0.8  | 0.6  |
| GeH <sub>5</sub> O <sup>+</sup>             | 0.2              | 0.1  | –    |
| GeH <sub>3</sub> O <sup>+</sup>             | 0.3              | 0.4  | 0.2  |
| GeH <sub>2</sub> O <sup>+</sup>             | 0.7              | 0.6  | 0.6  |
| GeHO <sup>+</sup>                           | 0.1              | 0.1  | 0.2  |
| GeH <sub>3</sub> C <sup>+</sup>             | 0.4              | 0.2  | –    |
| GeH <sub>2</sub> C <sup>+</sup>             | 0.5              | 0.4  | 0.2  |
| GeHC <sup>+</sup>                           | 0.3              | 0.2  | 0.2  |
| GeH <sub>3</sub> <sup>+</sup>               | 100              | 100  | 100  |
| GeH <sub>2</sub> <sup>+</sup>               | 25               | 42   | 64   |
| GeH <sup>+</sup>                            | 12               | 17   | 63   |
| Ge <sup>+</sup>                             | 14               | 22   | 88   |

<sup>a</sup> Abundances refer to <sup>76</sup>Ge isotope.

In high pressure mass spectrometry the same primary steps occur as in FTMS experiments: formation of CO<sup>+</sup> and GeH<sub>*n*</sub><sup>+</sup> (*n* = 1–3) ions and charge exchange at each collision between CO<sup>+</sup> and GeH<sub>4</sub> to give secondary GeH<sub>*n*</sub><sup>+</sup> (*n* = 0–3) fragments. The relative abundances of the ions in the mass spectra of a GeH<sub>4</sub>(9.83%)/CO mixture are reported in Table 1. The GeH<sub>*n*</sub><sup>+</sup> (*n* = 0–3) ions are the most abundant species and are involved in self-condensation reactions to give the Ge<sub>2</sub>H<sub>*n*</sub><sup>+</sup> (*n* = 1–5) dimers. Germanium-containing ions take part in ion-molecule reactions with CO and form three different ionic families: (i) GeH<sub>*n*</sub>CO<sup>+</sup> (*n* = 0–3) ions in which the entire molecule of the reagent gas is retained; (ii) GeH<sub>*n*</sub>O<sup>+</sup> (*n* = 1–3, 5) and (iii) GeH<sub>*n*</sub>C<sup>+</sup> (*n* = 0–3) ions in which one of the atoms of the reagent gas is lost in the neutral fragment. From the ion abundances it can be seen that the gas phase reactivity of GeH<sub>*n*</sub><sup>+</sup> (*n* = 0–3) and Ge<sub>2</sub>H<sub>*n*</sub><sup>+</sup> (*n* = 1–5) with carbon monoxide is very low, being 97.1 and 99.2% of the total current transported by GeH<sub>*n*</sub><sup>+</sup> (*n* = 0–3) in the GeH<sub>4</sub>(9.83%)/CO mixture at 0.3 and 0.03 torr, respectively. As the total pressure is decreased the relative abundances of oxygen- and carbon-containing ions also gradually decrease.

In the GeH<sub>4</sub>(4.0%)/CO mixture (Table 2) the same ions are observed. The Ge<sub>2</sub>H<sub>*n*</sub><sup>+</sup> (*n* = 1–5) dimers are weaker, whilst the GeH<sub>*n*</sub>CO<sup>+</sup>, GeH<sub>*n*</sub>O<sup>+</sup>, and GeH<sub>*n*</sub>C<sup>+</sup> ions become gradually stronger. These changes are due to the different relative concentrations of the reagents. A lower number of collisions between germanium-containing ionic and neutral species and a higher number of collisions with CO occur when the concentration of monogermane is decreased with respect to that of carbon monoxide. Under drastic experimental conditions, such as total pressure > 1

Table 2

Relative abundances of significant ions in the mass spectra of GeH<sub>4</sub>(4.0%)/CO mixtures at various pressures<sup>a</sup>

| Ions  | Pressures (torr) |      |      |      |
|---|------------------|------|------|------|
|   | 0.3              | 0.1  | 0.03 | 0.01 |
| Ge <sub>2</sub> H <sub>5</sub> <sup>+</sup> | 0.02             | 0.02 | –    | –    |
| Ge <sub>2</sub> H <sub>4</sub> <sup>+</sup> | 0.4              | 0.3  | 0.1  | 0.04 |
| Ge <sub>2</sub> H <sub>3</sub> <sup>+</sup> | 0.1              | 0.1  | 0.1  | 0.1  |
| Ge <sub>2</sub> H <sub>2</sub> <sup>+</sup> | 0.3              | 0.2  | –    | –    |
| Ge <sub>2</sub> H <sup>+</sup>              | 0.1              | –    | –    | –    |
| GeH <sub>3</sub> CO <sup>+</sup>            | 0.2              | 0.04 | –    | –    |
| GeH <sub>2</sub> CO <sup>+</sup>            | 0.2              | 0.1  | 0.2  | –    |
| GeHCO <sup>+</sup>                          | 0.5              | 0.5  | 0.3  | 0.1  |
| GeCO <sup>+</sup>                           | 1.3              | 1.4  | 1.2  | 0.8  |
| GeH <sub>5</sub> O <sup>+</sup>             | 0.1              | 0.02 | –    | –    |
| GeH <sub>3</sub> O <sup>+</sup>             | 0.4              | 0.6  | 0.4  | 0.4  |
| GeH <sub>2</sub> O <sup>+</sup>             | 0.6              | 1.1  | 0.9  | 0.9  |
| GeHO <sup>+</sup>                           | 0.05             | 0.1  | –    | –    |
| GeH <sub>3</sub> C <sup>+</sup>             | 0.2              | 0.1  | 0.2  | –    |
| GeH <sub>2</sub> C <sup>+</sup>             | 0.9              | 0.5  | 0.3  | 0.2  |
| GeHC <sup>+</sup>                           | 0.6              | 0.4  | 0.2  | –    |
| GeC <sup>+</sup>                            | 0.05             | 0.03 | 0.1  | –    |
| GeH <sub>3</sub> <sup>+</sup>               | 100              | 100  | 100  | 100  |
| GeH <sub>2</sub> <sup>+</sup>               | 39               | 53   | 69   | 93   |
| GeH <sup>+</sup>                            | 18               | 20   | 23   | 29   |
| Ge <sup>+</sup>                             | 20               | 26   | 33   | 44   |

<sup>a</sup> Abundances refer to <sup>76</sup>Ge isotope.

torr and with GeH<sub>4</sub>/CO ratio of about 1/10, the relative abundance of the GeH<sub>*n*</sub>C<sup>+</sup> (*n* = 0–3) ionic family strongly increases, and about 5% of the total ion current is transported by these carbon-containing ions.

The overall behaviour of the GeH<sub>4</sub>/CO mixtures is similar to that of the GeH<sub>4</sub>/O<sub>2</sub> mixtures under similar conditions [8]. In the presence of molecular oxygen, formation is observed of GeH<sub>*n*</sub>O<sub>2</sub><sup>+</sup> and GeH<sub>*n*</sub>O<sup>+</sup> ions containing the entire molecule of the reagent gas and one of its atoms, respectively. The corresponding GeH<sub>*n*</sub>CO<sup>+</sup>, GeH<sub>*n*</sub>O<sup>+</sup>, and GeH<sub>*n*</sub>C<sup>+</sup> ions originate in ion-molecule reactions with carbon monoxide. The main difference between the two systems consists in the much lower abundance of the oxygen- and carbon-containing species of the GeH<sub>4</sub>/CO mixtures.

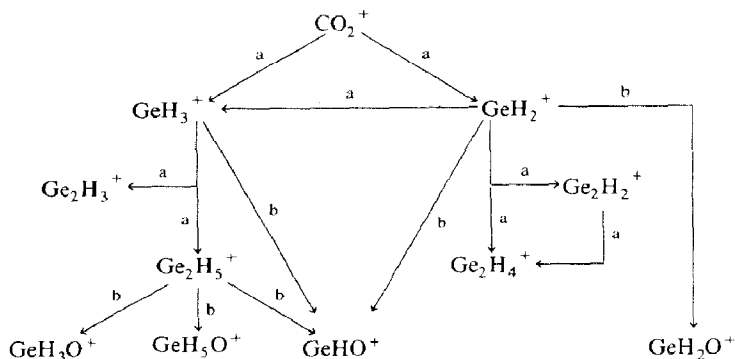
Of all these last ionic families only the GeH<sub>*n*</sub>O<sup>+</sup> species are detected by FTMS, similarly to what has been reported for the GeH<sub>4</sub>/O<sub>2</sub> mixtures [8]. As has been already pointed out, results from FTMS and CIMS experiments can be compared as far as bimolecular reactions are concerned [13]. Thus, the total number of collisions is about the same in the FTMS operating at 10<sup>–7</sup> torr and 1 s delay time and in the CIMS experiments at 10<sup>–2</sup> torr and 10<sup>–5</sup> delay time. In contrast, thermalization of ionic species which have an excess of internal energy takes place through two different processes depending on the collision frequency, and so it is likely that ionic species observed in the low pressure system (FTMS) after a delay of ca. 0.4 s are formed by unimolecular decomposition of excited ions, which in a high pressure

instrument (CIMS) are stabilized by unreactive collisions with the neutrals present in the ion chamber.

The data obtained in the high pressure experiments seem to be of interest mainly in respect of the presence of the  $\text{GeH}_n\text{C}^+$  ( $n = 1-3$ ) ions. By analogy with the hypothesis advanced previously [8] for the formation of  $\text{GeHO}^+$  from  $\text{Ge}_2\text{H}_2^+$  with  $\text{O}_2$ , it can be suggested that the  $\text{GeH}_n\text{C}^+$  species are formed by a multistep process in which a cyclic four-membered transition state involving carbon, oxygen, and two germanium atoms is involved, and, indeed,  $\text{GeH}_n\text{CO}^+$  ( $n = 0, 1, 3$ ) ions are observed. The successive cleavage of the carbon–oxygen and germanium–germanium bonds, with distribution of the charge between the two fragments, each containing an oxygen or a carbon atom, leads to  $\text{GeH}_n\text{O}^+$  ( $n = 1-3$ ) and  $\text{GeH}_n\text{C}^+$  ( $n = 1-3$ ) in parallel pathways.

#### *Ion-molecule reactions in $\text{GeH}_4/\text{CO}_2$ mixtures*

An ion pattern similar to that operating in  $\text{GeH}_4/\text{CO}$  mixtures is generated by 20 eV electron impact on  $\text{GeH}_4/\text{CO}_2$  (1/25) mixture at  $5 \times 10^{-7}$  torr total pressure under FTMS conditions. The main reactions, as monitored by double resonance experiments, are depicted in Scheme 2. Only ionic reagent/products are reported, the neutral reagent molecules being  $\text{GeH}_4$  (a) or  $\text{CO}_2$  (b). In addition to very minor abundances of the primary germane ions,  $\text{GeH}_n^+$  ( $n = 0-3$ ), electron bombardment produces essentially the  $\text{CO}_2^+$  (IP( $\text{CO}_2$ )  $13.769 \pm 0.03$  eV) [11] ions, which rapidly react with  $\text{GeH}_4$  to give predominantly  $\text{GeH}_3^+$  (ca. 75%) and  $\text{GeH}_2^+$  (ca. 25%),  $\text{GeH}_3^+$  being formed also by  $\text{GeH}_2^+$  as previously described. The primary and secondary  $\text{GeH}_n^+$  ( $n = 0-3$ ) fragments add to  $\text{GeH}_4$  yielding the normal digermene ion pattern. As in the case of  $\text{CO}$ , progressive loss of hydrogen from  $\text{Ge}_2\text{H}_n^+$  ( $n = 2-5$ ) by collisions with  $\text{CO}_2$  is observed. The correspondence between the  $\text{GeH}_4/\text{CO}$  and the  $\text{GeH}_4/\text{CO}_2$  ion-molecule reaction networks also extends to the mechanism of formation of the  $\text{GeH}_n\text{O}^+$  ( $n = 1, 3, 5$ ) ion family.  $\text{Ge}_2\text{H}_5^+$  is found to react with  $\text{CO}_2$  in about 100 ms to give  $\text{GeH}_5\text{O}^+$  (ca. 35%),  $\text{GeH}_3\text{O}^+$  (ca. 15%), and  $\text{GeHO}^+$  (ca. 50%). Further contribution to the  $\text{GeHO}^+$  ion abundance arises from comparatively slow condensation of  $\text{GeH}_n^+$  ( $n = 2, 3$ ) ions with a  $\text{CO}_2$  molecule. Moreover, attack of  $\text{GeH}_2^+$  on  $\text{CO}_2$  yields minor amounts of  $\text{GeH}_2\text{O}^+$ . By analogy with the conclusion reached for the  $\text{GeH}_4/\text{CO}$  mixtures, addition of  $\text{Ge}_2\text{H}_5^+$  to  $\text{CO}_2$  to yield  $\text{GeH}_n\text{O}^+$  ( $n = 1, 3, 5$ ) follows a complex reaction pathway



Scheme 2

involving formation of a Ge–O bond and extensive reshuffling of the hydrogen atoms in the addition complex.

Under chemical ionization conditions also the  $\text{GeH}_4/\text{CO}_2$  mixtures show a very similar behaviour to that displayed by the  $\text{GeH}_4/\text{CO}$  system. Most of the charge is transported by  $\text{GeH}_n^+$  ( $n = 0-3$ ) ions, which are mainly originated by charge-exchange processes between  $\text{CO}_2^+$  and  $\text{GeH}_4$ . In addition to the dimer species, the three ionic families of formula  $\text{GeH}_n\text{CO}_2^+$  ( $n = 0, 1$ ),  $\text{GeH}_n\text{O}^+$  ( $n = 0-3, 5$ ), and  $\text{GeH}_n\text{C}^+$  ( $n = 1-3$ ) are observed, as shown in Tables 3 and 4 for the  $\text{GeH}_4(10.5\%)/\text{CO}_2$  and the  $\text{GeH}_4(1.78\%)/\text{CO}_2$  mixtures, respectively. As it does towards CO,  $\text{GeH}_4$  displays a very low reactivity towards  $\text{CO}_2$  and, in the  $\text{GeH}_4(10.5\%)/\text{CO}_2$  mixture, the  $\text{GeH}_n\text{C}^+$  ions are barely detectable. Moreover, the changes of the ionic abundances with the variation of the total pressure follow the same trend with both CO and  $\text{CO}_2$ , and depend on the probability of collisions between germanium-containing ions and the reagent gas molecules. Once again, differences in relative abundances of ionic species under FTMS conditions compared with those under CIMS conditions can be ascribed to the more extensive decomposition processes which occur in the FTMS cell, and which lead, for example, to more abundant  $\text{GeHO}^+$  ions in FTMS experiments than in the high pressure mass spectra.

#### *Ion-molecule reactions in $\text{GeH}_4/\text{C}_2\text{H}_4$ mixtures*

Both  $\text{C}_2\text{H}_4^+$  and the germane ion family are observed after 12 eV electron bombardment of a  $\text{GeH}_4/\text{C}_2\text{H}_4$  mixture (1/10) at a pressure of  $5 \times 10^{-7}$  torr in the FTMS cell. The main reactions, as monitored by double resonance experiments, are shown in Scheme 3. Only ionic reagents/products are reported, the neutral reagent molecules being  $\text{GeH}_4$  (a) or  $\text{C}_2\text{H}_4$  (b). Reactive collisions between  $\text{C}_2\text{H}_4^+$  and the neutrals present lead to its rapid disappearance with formation of  $\text{C}_3\text{H}_5^+$

Table 3

Relative abundances of significant ions in the mass spectra of  $\text{GeH}_4(10.5\%)/\text{CO}_2$  mixtures at various pressures <sup>a</sup>

| Ions                      | Pressures (torr) |      |      |      |
|---------------------------|------------------|------|------|------|
|                           | 0.4              | 0.1  | 0.03 | 0.01 |
| $\text{Ge}_2\text{H}_5^+$ | 0.1              | –    | –    | –    |
| $\text{Ge}_2\text{H}_4^+$ | 0.2              | 0.1  | 0.1  | 0.01 |
| $\text{Ge}_2\text{H}_3^+$ | 0.1              | 0.05 | 0.02 | 0.01 |
| $\text{Ge}_2\text{H}_2^+$ | 0.2              | 0.1  | 0.1  | 0.1  |
| $\text{Ge}_2\text{H}^+$   | 0.01             | –    | –    | –    |
| $\text{GeCO}_2^+$         | 0.1              | 0.1  | 0.1  | 0.1  |
| $\text{GeH}_5\text{O}^+$  | 0.2              | –    | 0.1  | 0.3  |
| $\text{GeH}_3\text{O}^+$  | 1.6              | 1.1  | 0.9  | 0.7  |
| $\text{GeH}_2\text{O}^+$  | 0.9              | 0.8  | 0.7  | 0.5  |
| $\text{GeHO}^+$           | 1.2              | 0.9  | 0.6  | 0.5  |
| $\text{GeH}_3\text{C}^+$  | –                | 0.1  | 0.1  | –    |
| $\text{GeH}_2\text{C}^+$  | 0.1              | 0.1  | –    | –    |
| $\text{GeH}_3^+$          | 100              | 100  | 100  | 100  |
| $\text{GeH}_2^+$          | 59               | 74   | 88   | 95   |
| $\text{GeH}^+$            | 12               | 17   | 22   | 22   |
| $\text{Ge}^+$             | 21               | 28   | 36   | 36   |

<sup>a</sup> Abundances refer to <sup>76</sup>Ge isotope.

Table 4

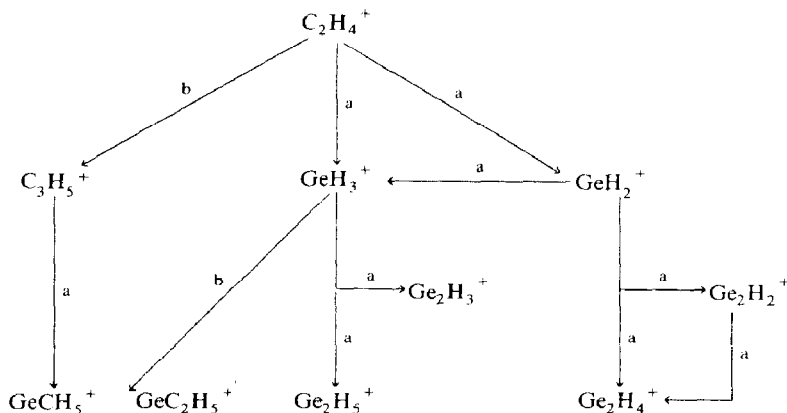
Relative abundances of significant ions in the mass spectra of GeH<sub>4</sub>(1.78%)/CO<sub>2</sub> mixtures at various pressures<sup>a</sup>

| Ions  | Pressures (torr) |     |      |
|---|------------------|-----|------|
|   | 0.4              | 0.1 | 0.03 |
| Ge <sub>2</sub> H <sub>5</sub> <sup>+</sup> | –                | 0.2 | –    |
| Ge <sub>2</sub> H <sub>4</sub> <sup>+</sup> | 0.4              | 0.2 | 0.1  |
| Ge <sub>2</sub> H <sub>3</sub> <sup>+</sup> | 0.1              | 0.1 | –    |
| Ge <sub>2</sub> H <sub>2</sub> <sup>+</sup> | 0.3              | 0.3 | 0.2  |
| GeHCO <sub>2</sub> <sup>+</sup>             | 0.2              | 0.3 | 0.3  |
| GeCO <sub>2</sub> <sup>+</sup>              | 1.1              | 1.4 | 1.1  |
| GeH <sub>5</sub> O <sup>+</sup>             | 0.6              | 0.5 | 0.3  |
| GeH <sub>3</sub> O <sup>+</sup>             | 0.5              | 1.2 | 1.0  |
| GeH <sub>2</sub> O <sup>+</sup>             | 1.4              | 0.4 | –    |
| GeHO <sup>+</sup>                           | –                | 0.1 | –    |
| GeO <sup>+</sup>                            | 1.0              | 1.5 | 1.3  |
| GeH <sub>3</sub> C <sup>+</sup>             | 1.4              | 0.4 | –    |
| GeH <sub>2</sub> C <sup>+</sup>             | 0.2              | –   | 0.3  |
| GeHC <sup>+</sup>                           | 0.2              | 0.4 | 0.6  |
| GeH <sub>3</sub> <sup>+</sup>               | 100              | 100 | 100  |
| GeH <sub>2</sub> <sup>+</sup>               | 48               | 70  | 88   |
| GeH <sup>+</sup>                            | 8                | 12  | 21   |
| Ge <sup>+</sup>                             | 15               | 18  | 31   |

<sup>a</sup> Abundances refer to <sup>76</sup>Ge isotope.

from C<sub>2</sub>H<sub>4</sub> and of GeH<sub>3</sub><sup>+</sup> (with very minor yield of GeH<sub>2</sub><sup>+</sup>) from GeH<sub>4</sub>. The GeH<sub>*n*</sub><sup>+</sup> (*n* = 0–3) fragments react with GeH<sub>4</sub> to produce the digermene daughter species, Ge<sub>2</sub>H<sub>*n*</sub><sup>+</sup> (*n* = 2–5). The neutral C<sub>2</sub>H<sub>4</sub> effectively competes with GeH<sub>4</sub> for GeH<sub>3</sub><sup>+</sup>, yielding GeC<sub>2</sub>H<sub>5</sub><sup>+</sup>. Similarly effective is the attack of C<sub>3</sub>H<sub>5</sub><sup>+</sup> on GeH<sub>4</sub> to give GeCH<sub>5</sub><sup>+</sup>. Under the conditions prevailing in the FTMS experiments, the Ge<sub>2</sub>H<sub>*n*</sub><sup>+</sup> (*n* = 2–5) ion family is found not to react with C<sub>2</sub>H<sub>4</sub> even at relatively long reaction times (i.e. 400 ms). It is noteworthy that in this system germanium-containing analogues of the well known carbocations C<sub>2</sub>H<sub>5</sub><sup>+</sup> (i.e. GeCH<sub>5</sub><sup>+</sup>) and C<sub>3</sub>H<sub>5</sub><sup>+</sup> (i.e. GeC<sub>2</sub>H<sub>5</sub><sup>+</sup>) are observed.

Under high pressure conditions, abundant ions containing germanium–carbon bonds are observed in the mass spectra of the GeH<sub>4</sub>(10.1%)/C<sub>2</sub>H<sub>4</sub> mixture, and the



Scheme 3



Table 5

Relative abundances of significant ions in the mass spectra of GeH<sub>4</sub>(10.1%)/C<sub>2</sub>H<sub>4</sub> mixtures at various pressures <sup>a</sup>

| Ions  | Pressures (torr) |     |      |      |
|---|------------------|-----|------|------|
|   | 0.5              | 0.2 | 0.05 | 0.01 |
| GeC <sub>4</sub> H <sub>11</sub> <sup>+</sup> | 4.8              | 2.0 | –    | –    |
| GeC <sub>4</sub> H <sub>10</sub> <sup>+</sup> | 1.0              | 0.4 | –    | –    |
| GeC <sub>4</sub> H <sub>9</sub> <sup>+</sup>  | 1.0              | 0.3 | –    | –    |
| GeC <sub>4</sub> H <sub>8</sub> <sup>+</sup>  | 0.3              | 0.3 | –    | –    |
| GeC <sub>4</sub> H <sub>7</sub> <sup>+</sup>  | 0.3              | 0.5 | –    | –    |
| GeC <sub>3</sub> H <sub>9</sub> <sup>+</sup>  | 9.9              | 6.2 | 0.5  | –    |
| GeC <sub>3</sub> H <sub>8</sub> <sup>+</sup>  | 2.1              | 1.0 | –    | –    |
| GeC <sub>3</sub> H <sub>5</sub> <sup>+</sup>  | 0.3              | 0.2 | –    | –    |
| GeC <sub>2</sub> H <sub>7</sub> <sup>+</sup>  | 16               | 11  | 3.2  | 0.7  |
| GeC <sub>2</sub> H <sub>6</sub> <sup>+</sup>  | 3.0              | 2.3 | –    | –    |
| GeC <sub>2</sub> H <sub>5</sub> <sup>+</sup>  | 55               | 47  | 15   | 8.0  |
| GeC <sub>2</sub> H <sub>4</sub> <sup>+</sup>  | 10               | 8.4 | 5.3  | 3.3  |
| GeC <sub>2</sub> H <sub>3</sub> <sup>+</sup>  | 3.2              | 3.1 | 2.4  | 1.5  |
| GeCH <sub>5</sub> <sup>+</sup>                | 22               | 22  | 12   | 5.0  |
| GeCH <sub>4</sub> <sup>+</sup>                | 3.2              | 2.8 | 3.0  | 2.4  |
| GeCH <sub>3</sub> <sup>+</sup>                | 5.2              | 3.2 | 2.1  | 1.1  |
| GeCH <sub>2</sub> <sup>+</sup>                | 3.6              | 0.8 | –    | 0.2  |
| GeH <sub>3</sub> <sup>+</sup>                 | 100              | 100 | 100  | 100  |
| GeH <sub>2</sub> <sup>+</sup>                 | 11               | 14  | 31   | 59   |
| GeH <sup>+</sup>                              | 10               | 16  | 32   | 23   |
| Ge <sup>+</sup>                               | 60               | 65  | 67   | 65   |

<sup>a</sup> Abundances refer to <sup>76</sup>Ge isotope.

Table 6

Relative abundances of significant ions in the mass spectra of GeH<sub>4</sub>(1.0%)/C<sub>2</sub>H<sub>4</sub> mixtures at various pressures <sup>a</sup>

| Ions  | Pressures (torr) |     |      |      |
|---|------------------|-----|------|------|
|   | 0.5              | 0.2 | 0.05 | 0.01 |
| GeC <sub>4</sub> H <sub>11</sub> <sup>+</sup> | 0.7              | 0.4 | –    | –    |
| GeC <sub>4</sub> H <sub>10</sub> <sup>+</sup> | 0.1              | –   | –    | –    |
| GeC <sub>4</sub> H <sub>9</sub> <sup>+</sup>  | 0.3              | 0.3 | –    | –    |
| GeC <sub>4</sub> H <sub>8</sub> <sup>+</sup>  | 0.1              | –   | –    | –    |
| GeC <sub>3</sub> H <sub>9</sub> <sup>+</sup>  | 2.0              | 1.4 | 0.3  | –    |
| GeC <sub>3</sub> H <sub>8</sub> <sup>+</sup>  | 0.5              | 0.2 | –    | –    |
| GeC <sub>3</sub> H <sub>5</sub> <sup>+</sup>  | 0.1              | –   | –    | –    |
| GeC <sub>2</sub> H <sub>7</sub> <sup>+</sup>  | 2.6              | 2.1 | 1.3  | –    |
| GeC <sub>2</sub> H <sub>6</sub> <sup>+</sup>  | 0.8              | 0.6 | –    | –    |
| GeC <sub>2</sub> H <sub>5</sub> <sup>+</sup>  | 10               | 11  | 12   | 6.5  |
| GeC <sub>2</sub> H <sub>4</sub> <sup>+</sup>  | 2.5              | 2.1 | 2.0  | –    |
| GeC <sub>2</sub> H <sub>3</sub> <sup>+</sup>  | 1.1              | 0.8 | 0.9  | 1.5  |
| GeCH <sub>5</sub> <sup>+</sup>                | 4.3              | 5.0 | 2.5  | 1.0  |
| GeCH <sub>4</sub> <sup>+</sup>                | 1.7              | 1.6 | 1.4  | 0.7  |
| GeCH <sub>3</sub> <sup>+</sup>                | 1.7              | 1.4 | 1.4  | 1.5  |
| GeCH <sub>2</sub> <sup>+</sup>                | 2.1              | 1.3 | 0.7  | 0.9  |
| GeH <sub>3</sub> <sup>+</sup>                 | 100              | 100 | 100  | 100  |
| GeH <sub>2</sub> <sup>+</sup>                 | 7.8              | 9.5 | 22   | 48   |
| GeH <sup>+</sup>                              | 9.0              | 9.8 | 14   | 23   |
| Ge <sup>+</sup>                               | 16               | 20  | 46   | 84   |

<sup>a</sup> Abundances refer to <sup>76</sup>Ge isotope.

monogermanium ones are reported in Table 5. The highest mass ions are  $\text{GeC}_4\text{H}_n^+$  ( $n = 7-11$ ), but species containing three, two, or one carbon atoms are also formed at 0.5, 0.2 and 0.05 torr total pressure. The base peak is again  $\text{GeH}_3^+$  in this system at all pressures investigated, but at 0.5 torr only 56% of the total charge is carried by  $\text{GeH}_n^+$  ( $n = 0-3$ ), while 27.2 and 10.7% of the total charge is carried by  $\text{GeC}_2\text{H}_n^+$  ( $n = 3-7$ ) and by  $\text{GeCH}_n^+$  ( $n = 2-5$ ), respectively.

When ethylene alone is introduced into the ion source at about 0.5 torr,  $\text{C}_2\text{H}_4^+$  is the base peak and abundant ions such as  $\text{C}_2\text{H}_5^+$ ,  $\text{C}_3\text{H}_3^+$ ,  $\text{C}_3\text{H}_5^+$  (second most abundant ion),  $\text{C}_4\text{H}_5^+$ ,  $\text{C}_4\text{H}_7^+$ ,  $\text{C}_5\text{H}_7^+$ , and  $\text{C}_5\text{H}_9^+$  are formed by ion-molecule reactions. Moreover, a decreasing total pressure corresponds to a gradual disappearance of the highest mass ions, which parallels the disappearance of the hydrocarbon ions containing five and four carbon atoms, and a decrease of the abundances of the lightest ionic species. Therefore, it is likely that formation of most of the ions observed in the mass spectra of  $\text{GeH}_4/\text{C}_2\text{H}_4$  mixtures at high pressure occurs by reactions between neutral  $\text{GeH}_4$  and carbon-containing ions originated from  $\text{C}_2\text{H}_4$ . This formation process is also in agreement with data of Table 6, where the abundances of monogermanium ions observed in the  $\text{GeH}_4(1.0\%)/\text{C}_2\text{H}_4$  mixture are reported. There is a strong decrease in the abundances of the  $\text{GeC}_n\text{H}_m^+$  ( $n = 1-4$ ) ionic species when the concentration of  $\text{GeH}_4$  is reduced. The number of collisions between neutral  $\text{GeH}_4$  and carbon-containing ions from  $\text{C}_2\text{H}_4$ , whose abundances are about the same at the same pressure in the two different mixtures, is lower in the  $\text{GeH}_4(1.0\%)/\text{C}_2\text{H}_4$  mixture with respect to the  $\text{GeH}_4(10.1\%)/\text{C}_2\text{H}_4$  one, causing a reduced abundance of the reaction ionic products.

It is noteworthy that in both mixtures and at the highest pressures used,  $\text{Ge}_2\text{H}_n^+$  ( $n = 2-5$ ),  $\text{Ge}_2\text{C}_2\text{H}_5^+$ , and  $\text{Ge}_2\text{C}_4\text{H}_8^+$  ions are also observed, though they are generally much weaker than the corresponding monogermanium species.

## Conclusions

From these results interesting insights on the germanium chemistry can be obtained in respect of fundamental chemistry and the possibility of applications. The formation in the  $\text{GeH}_4/\text{CO}$  and  $\text{GeH}_4/\text{CO}_2$  systems of ions such as  $\text{GeH}_n\text{O}^+$ , previously observed in the presence of molecular oxygen, and  $\text{GeH}_n\text{C}^+$  is indicative of new reactions of  $\text{GeH}_4$ . Furthermore, the production of germanium-containing analogues of  $\text{C}_2\text{H}_5^+$  and  $\text{C}_3\text{H}_5^+$ , i.e.  $\text{GeCH}_5^+$  and  $\text{GeC}_2\text{H}_5^+$ , respectively, in the  $\text{GeH}_4/\text{C}_2\text{H}_4$  systems provides valuable information on the reactivity of germanium-hydrogen compounds in relation to that of hydrocarbons. The study of the structures and stability of all these ionic species will be further explored.

Carbon monoxide and carbon dioxide show a low reactivity towards monogermane. Moreover, reactions of  $\text{GeH}_4$  with CO and  $\text{CO}_2$  under any experimental conditions give oxygen-containing ionic species, whose presence is undesirable in amorphous solids to be used in photovoltaic devices. In contrast, the high abundances of the  $\text{GeC}_n\text{H}_m^+$  species (i.e. 44% of the total ion current in the  $\text{GeH}_4(10.1\%)/\text{C}_2\text{H}_4$  mixture at 0.5 torr total pressure), observed in the mixtures of  $\text{GeH}_4$  with  $\text{C}_2\text{H}_4$ , suggest that ethylene is a good reagent for the preparation of germanium carbides by decomposition of volatile hydrides. The photovoltaic properties of these carbides depend on the value of  $x$  in  $\text{Ge}_x\text{C}_{1-x}$ , and the results presented in this paper show that it can be adjusted by changing the total pressure,

the relative concentrations of the reagent gases, and the nature of the carbon-containing substrate.

## Experimental

GeH<sub>4</sub> was prepared by reduction of GeO<sub>2</sub> with KBH<sub>4</sub> and purified as described previously [7c].

Chemical ionization experiments were carried out with a double focussing Kratos MS80 mass spectrometer connected to a DS55 data system. Typical operating conditions were as follows: primary ionization electron energy 60 eV, emission current 500 μA, acceleration voltage 4 kV, ion repeller voltage zero, resolution 1000 (10% valley definition), scan rate 3 s/dec. The reaction mixtures containing various ratios of GeH<sub>4</sub> and reagent gas were prepared in a vial (1.50 l) connected to a vacuum line. It was filled with the appropriate amounts of the two gases to a total pressure of about 850 torr and connected to a valve of the multiple reagent gas inlet system of the mass spectrometer. The gaseous mixture was then introduced into the ion source, which was maintained at room temperature. Ion source pressures were evaluated from the relative concentrations of ions in CH<sub>4</sub> as a function of the pressure [14], and corrected by use of the relative sensitivities of the ion gauge for different gases [15]. However, small differences in pressure do not cause drastic variations in the nature and abundance of ions.

All FTMS experiments were performed with a Nicolet FTMS 1000 spectrometer equipped with a 2.54 cm cubic trapping cell situated between the poles of a superconducting magnet operated at 2.0 T. Typical operating conditions were as follows: nominal electron beam energy 12–20 eV, emission current 400 nA, electron beam pulse 3–5 ms, total pressure  $5 \times 10^{-7}$  torr, resolution 1000 fwhh (full width half height) at mass 100. Sample pressures (uncorrected) were measured with a Granville–Phyllips 280 Bayard–Alpert Ion Gauge. Inlet system and trapping cell were kept at room temperature. Ions were trapped in the cell by a differential trapping voltage of about 1.0 volt. The ionic sequences taking place in the reaction cell were examined by double resonance experiments by choosing suitable time sequences to maximize the abundance of a given precursor and suitable frequency windows to remove all the undesired ions from the cell. Detection of the daughter ion pattern after an appropriate delay time provides direct information about the reaction sequence concerning the precursor under investigation.

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